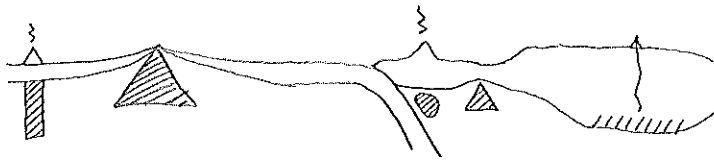


MELTING THE MANTLE

OBJECTIVES

$T, P, T_p, X, F, N_{0.8}$

1. MOTIVATION



IMPORTANT TO UNDERSTAND HOW + WHY MELTING OCCURS

- MELTING → CRUST FORMATION, PLANETARY DIFFERENTIATION
- MELTING → DEGASSING → ATMOSPHERE
- MELTING → REDISTRIBUTION OF TRACE ELEMENTS

MELTS CAN GIVE INSIGHT INTO

- THERMAL STATE OF EARTH'S INTERIOR
- COMPOSITION OF EARTH'S INTERIOR

KEY QUANTITIES WE WILL WANT ARE

T - temperature of melting

P - of melting

F - degree of melting (fraction of melting)

2. WHAT IS MELTING?

PHASE CHANGE FROM SOLID TO LIQUID

WHAT IS A SOLID ... LIQUID?

- SOLID MINERAL PHASE HAS A CRYSTALLINE STRUCTURE
LONG RANGE ORDER



LO ENTROPY

e.g. SiO_2 , WHERE ALL SiO_4^{4-} ARE BONDED

- LIQUID - SHORT RANGE ORDER



HI ENTROPY

e.g. molten SiO_2 ; some bonds are broken

* BE CAREFUL NOT TO CONFUSE THE DEFINITION OF A LIQUID WITH A RHEOLOGICAL FLUID.

3. MELTING OF A 1-COMPONENT (e.g. pure) SYSTEM

ENERGY NEED TO MELT

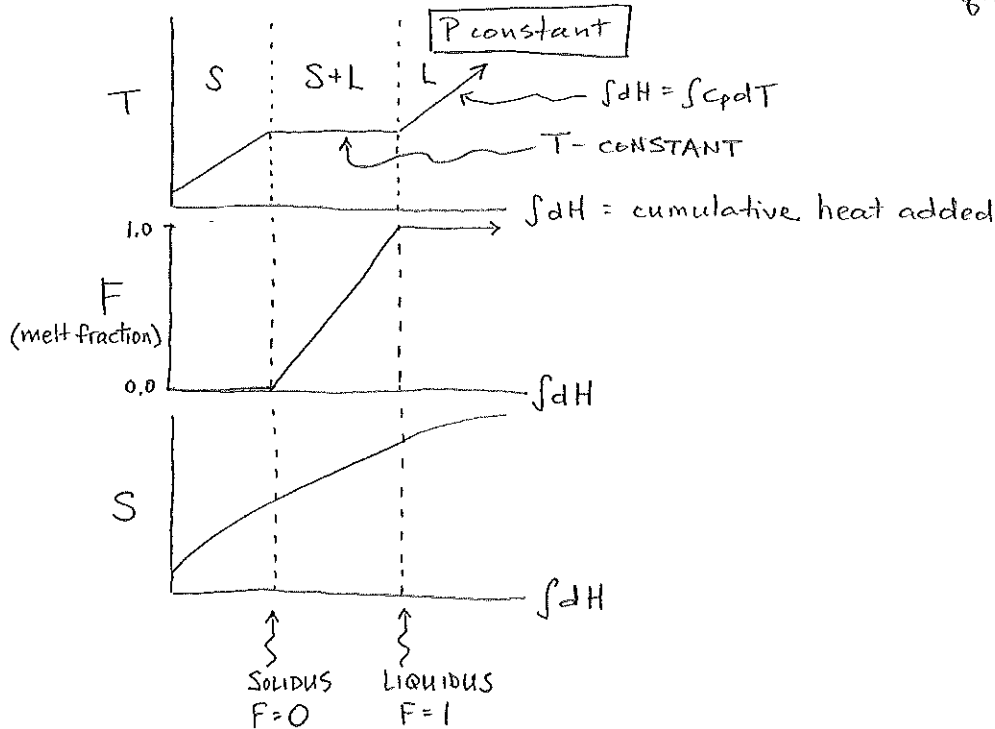
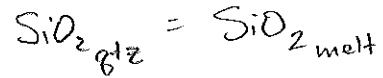
↳ HEAT OF FUSION ΔH_f

• THIS HEAT IS USED TO BREAK OR STRETCH BONDS

↳ ENTROPY INCREASES

$$\Delta S = \frac{\Delta H_f}{T_m}$$

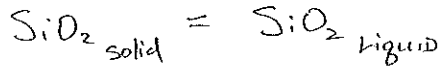
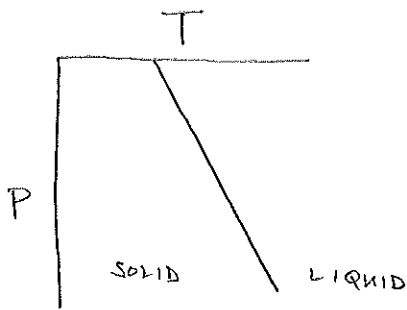
EXAMPLE,



* MELTING IS A PHASE CHANGE

* IN 1-COMPONENT SYSTEM, 2 PHASES COEXIST DURING MELTING
AT CONSTANT P, T CANNOT CHANGE UNTIL ONE PHASE IS CONSUMED.

MELTING PHASE DIAGRAM IN P-T SPACE - 1 COMPONENT SYSTEM



$$\Delta V_{\text{rxn}} = V_{\text{liq}} - V_{\text{solid}} > 0$$

$$\Delta S_{\text{rxn}} = S_{\text{liq}} - S_{\text{solid}} > 0$$

SEE BELOW ↓

$$\left(\frac{dP}{dT}\right)_{\text{rxn}} = \frac{\Delta S_{\text{rxn}}}{\Delta V_{\text{rxn}}} > 0$$

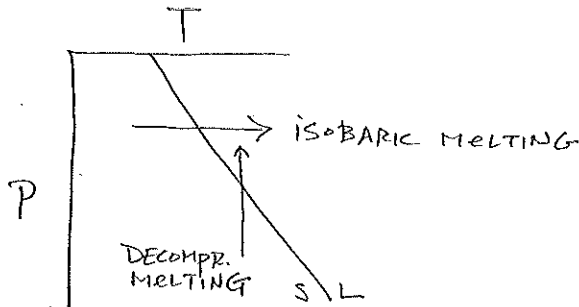
↑ CLAPEYRON SLOPE

IN GENERAL $\frac{dP}{dT} > 0$ FOR MELTING RXNS

NOTE $\left(\frac{dP}{dT}\right)_{\text{rxn}} = \frac{\Delta H_f}{T_f \Delta V_f}$
↑ fusion

IF > 0 , THEN ENDOTHERMIC (ASSUMING $\Delta V_f > 0$)

→ MELTING CAN OCCUR BY INCREASE IN T, DECREASE P



* NOTE: WITH INC. P $\Delta V_{\text{melting}} \rightarrow 0$



* DERIVATION OF CLAPEYRON SLOPE

1st LAW

internal energy = heat + work

$$dU = q + w$$

$$dU = TdS - PdV$$

$$U(S, V) = TS - PV$$

DEFINE G (USING LEGENDRE TRANSF.)

$$G \equiv U - TS + PV = H - TS$$

$$G \equiv H - TS \leftarrow \text{entropy}$$

Gibbs Free Energy ← enthalpy (heat + bond energy)

$$dG = dU - TdS - SdT + VdP - PdV$$

combine w/ $dU = TdS - PdV$

$$dG = -SdT + VdP$$

FOR A RXN $d\Delta G_{\text{rxn}} = -\Delta S_{\text{rxn}}dT + \Delta V_{\text{rxn}}dP$

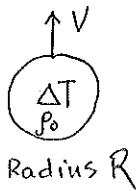
$$d\Delta G_{\text{rxn}} = 0 \text{ at equilibrium} \Rightarrow \frac{dP}{dT} = \frac{\Delta S_{\text{rxn}}}{\Delta V_{\text{rxn}}}$$

DECOMPRESSION MELTING IS IMPORTANT IN EARTH

↑ ΔP large, ΔT small

How DOES DECOMPRESSION MELTING OCCUR?

* CONSIDER A HOT BLOB IN THE MANTLE.



$\Delta \rho = \rho_0 \alpha \Delta T$ thermal expansion \rightarrow density decrease

IF BUOYANT FORCE EXCEEDS VISCOUS RESISTANCE...

UPWARDS VELOCITY $V \sim \frac{\Delta \rho g R^2}{\eta}$

BLOB WANTS TO COOL

COOLING TIME $\sim \frac{R^2}{K}$

CHARACTERISTIC TRANSPORT TIME $\sim \frac{R}{V}$

WHEN $\frac{\text{COOLING } t}{\text{ASCENT } t} \gg 1$, RISES FASTER THAN IT COOLS

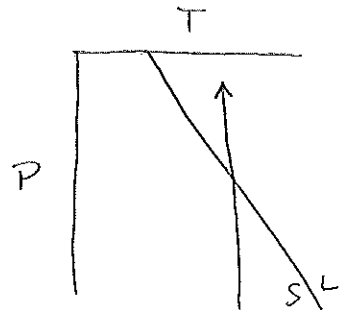
$$\frac{\Delta \rho g R^3}{\eta K} = \frac{\rho_0 \alpha g \Delta T R^3}{\eta K} = \text{Rayleigh number}$$

if $Ra > Ra_{cr} \sim 10^3$ CONVECTION

EARTH $Ra \sim 10^5 - 10^7$

CONCLUSION \rightarrow CONVECTION LEADS TO ADIABATIC UPWELLING

e.g. $q = 0$



ADIABATIC DECOMPRESSION

$$q=0$$

since $TdS \equiv q$, for a reversible process $\rightarrow dS=0$

ISENTROPIC PROCESS

$$\Delta U = q + W$$

$$\Delta U = TdS - PdV$$

isentropic

$$\Delta U = 0 - PdV$$

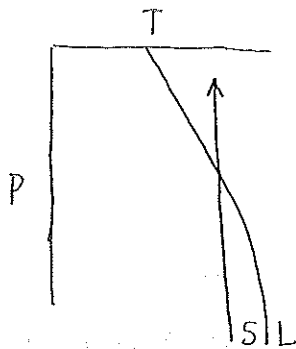
Isentropic decompression causes upwelling mantle to do work on surroundings, so internal energy drops (and T drops)

$$\left(\frac{\partial T}{\partial P}\right)_S = \frac{\alpha T}{\rho C_p}$$

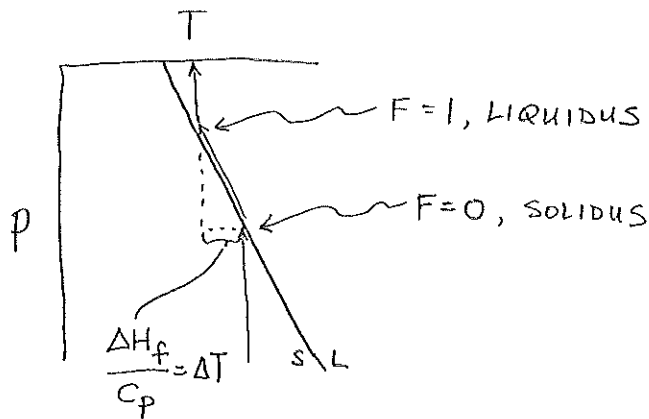
$$\text{OR } \left(\frac{\partial T}{\partial Z}\right)_S = \frac{\alpha g T}{C_p}$$

$$\text{SOLID } \left(\frac{\partial T}{\partial Z}\right)_S \sim 0.5^\circ \text{C/km}$$

Compare with SOLIDUS $\sim 3-4^\circ \text{C/km}$ (NOTE THIS VARIES WITH DEPTH)

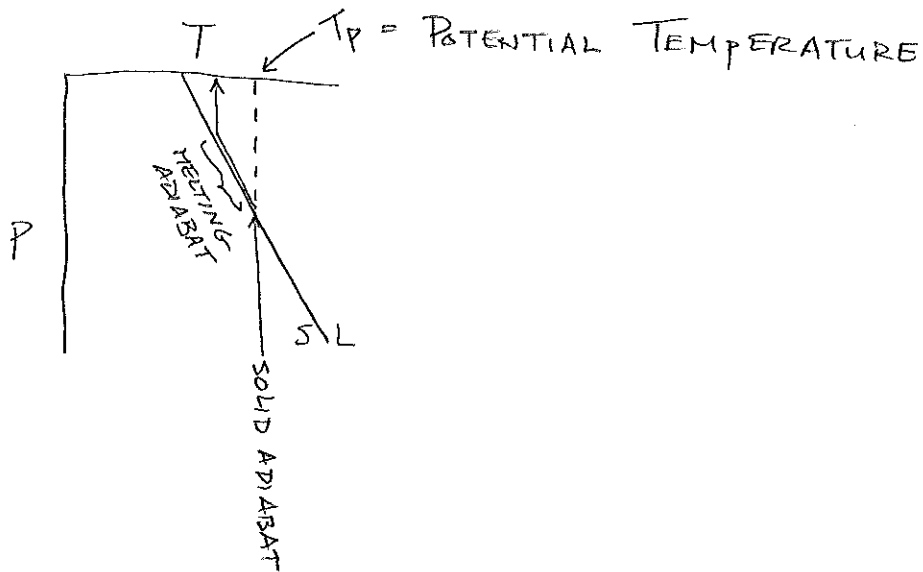


WHAT DOES ISENTROPIC MELTING LOOK LIKE IN A
1-COMPONENT SYSTEM



* AT CONSTANT P,
 $T_{\text{SOLIDUS}} = T_{\text{LIQUIDUS}}$
 * BUT NOT SO, DURING
 DECOMPRESSION

T DROPS DURING MELTING BECAUSE HEAT ABSORBED INTO
 PHASE CHANGE



MELTING IN MULTI-COMPONENT SYSTEM

BINARY EXAMPLE

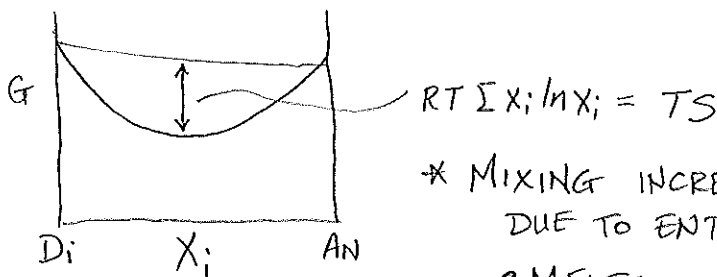
$$\bar{G}_{TIP} = \sum X_i \mu_i$$

\bar{G}_{TIP} : molar Gibbs free energy of phase
 X_i : mole fraction of component
 μ_i : chemical potential

$$\mu_i = \mu_i^0 + RT \ln X_i \quad (\text{assume ideal mixing})$$

$$\bar{G}_{TIP} = \sum X_i \mu_i^0 + RT \sum X_i \ln X_i$$

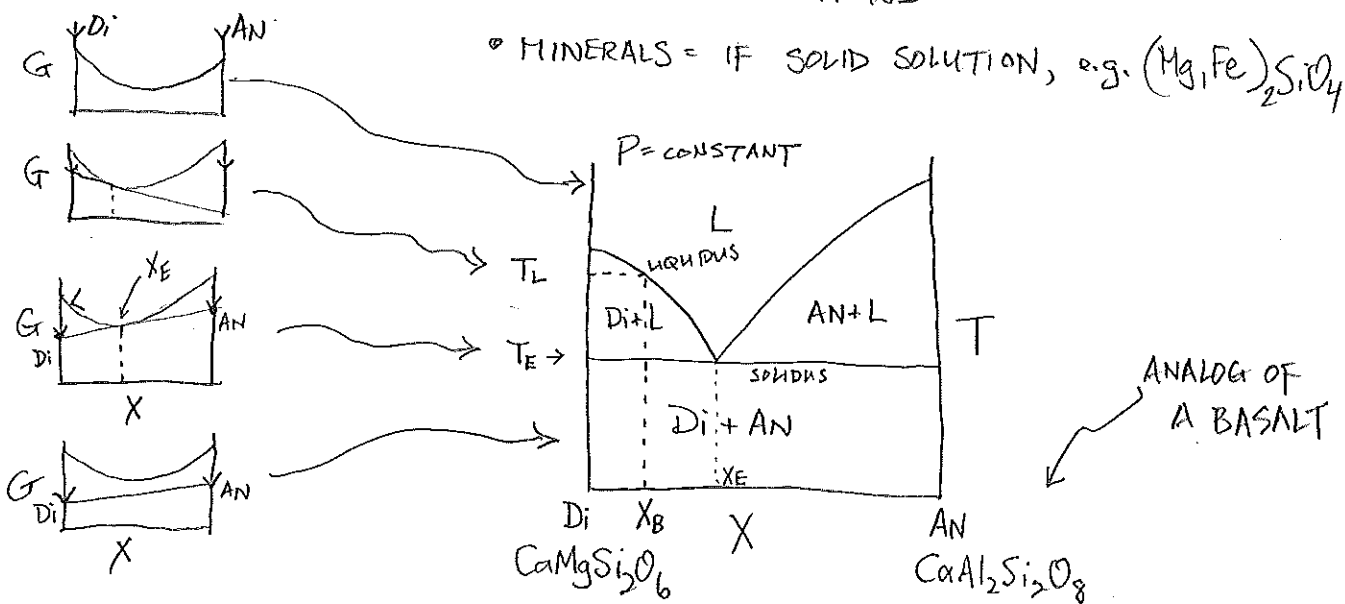
note similarity to $\Delta S_{mixing} = -R \sum X_i \ln X_i$



* MIXING INCREASES STABILITY OF A PHASE DUE TO ENTROPY INCREASE \rightarrow G DECREASE

• MELTS = SOLUTIONS

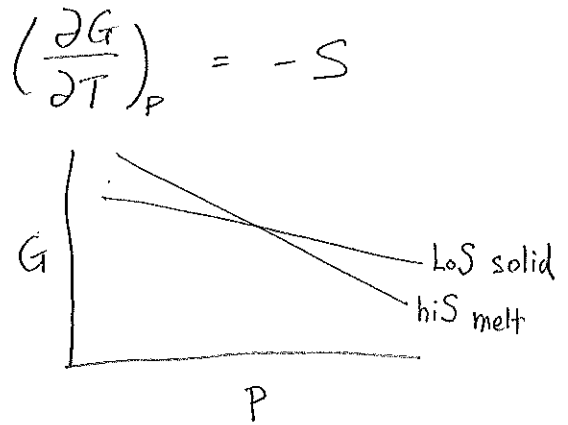
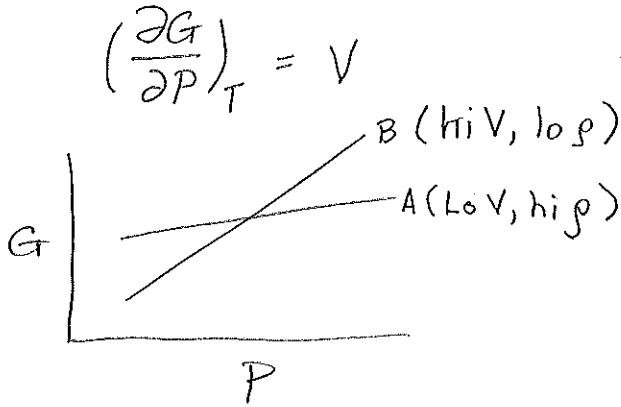
• MINERALS = IF SOLID SOLUTION, e.g. $(Mg, Fe)_2SiO_4$



RULES OF THUMB

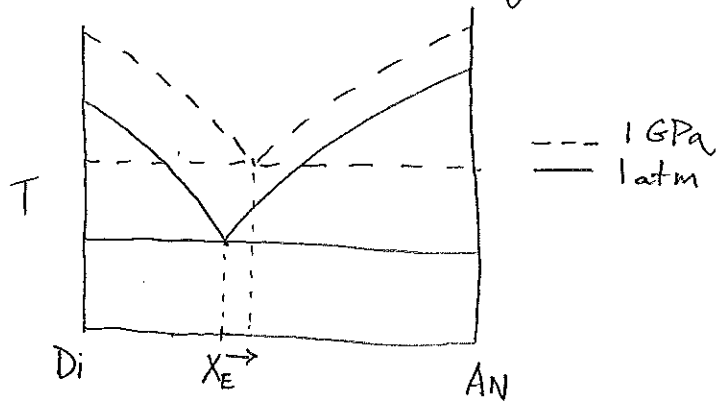
- ALL LIQUIDS AT THE ONSET OF MELTING HAVE SAME COMPOSITION (IN THIS, EUTECTIC COMP., X_E)
- $T_E = T_{SOLIDUS}$
- $T_{LIQUIDUS} \neq T_{SOLIDUS}$

BINARY XMP, EFFECT OF P

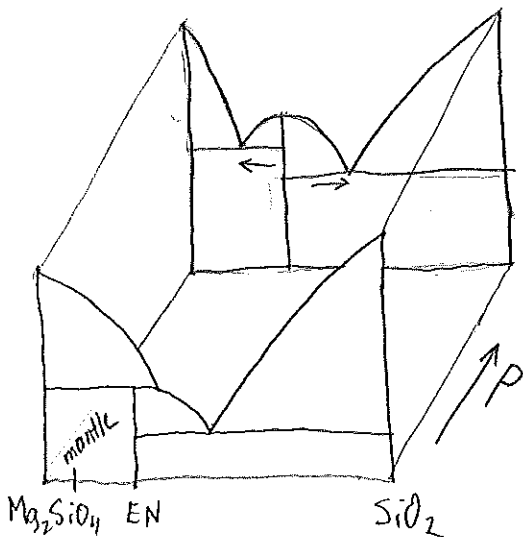


RULES OF THUMB

- * Hi P favors THE STABILITY OF DENSE PHASES
 ∴ HARDER TO MELT AT HIGHER P, ∴ T_E RISES
 ∴ X_E , EUTECTIC COMPOSITION, MOVES AWAY FROM hi P PHASE
 (i.e. phase field for hi P phase expands)
- * Hi T favors hi entropy phases
 melts require hi T.



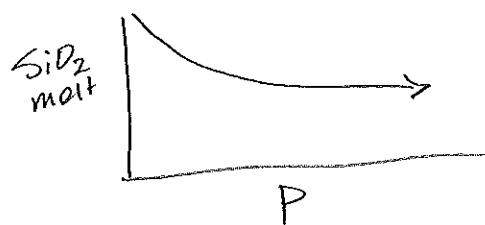
BASALT ANALOG



MANTLE PERIDOTITE ANALOG

Mg_2SiO_4 OLIVINE
 ENSTATITE $Mg_2Si_2O_6$

$P \uparrow$ SiO_2 in melt \downarrow



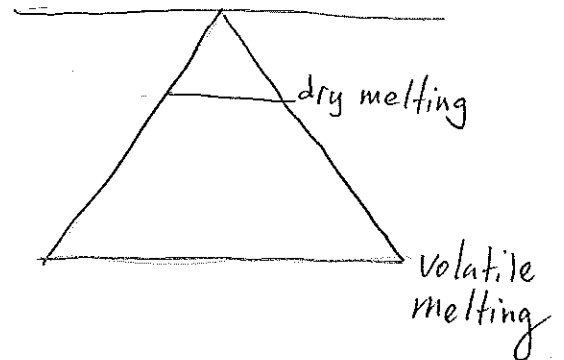
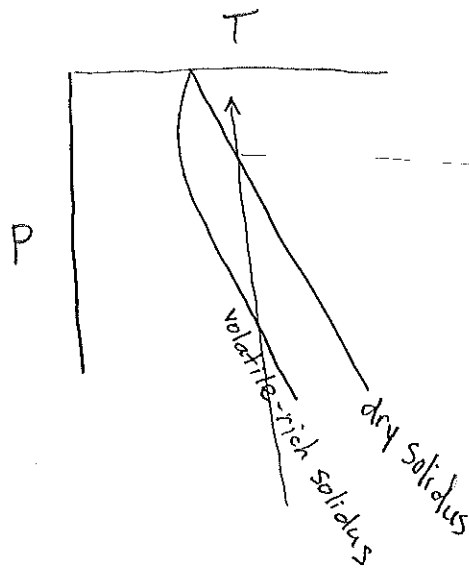
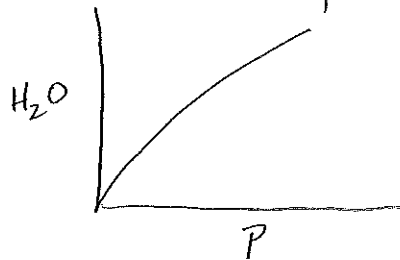
EFFECT OF VOLATILES

SOLIDUS IS DEPRESSED IF VOLATILES ARE SOLUBLE IN MELT.

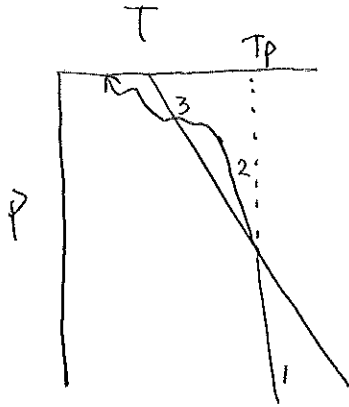
H_2O , CO_2 SOLUBLE IN SILICATE MELT

CH_4 , C NOT SOLUBLE SO REDOX MATTERS

BECAUSE SOLUBILITY VARIES AS P , FREEZING POINT DEPRESSION IS SMALLER AT LO P .



GOAL IS TO GET T_p BECAUSE THIS GIVES THERMAL STATE



- 1- solid adiabat
- 2- melting adiabat
- 3- non-adiabatic cooling

IF WE KNOW T , F OR P OF PRIMARY MAGMA THEN WE CAN SAY SOMETHING ABOUT T_p .

PRIMARY = IN EQUILIBRIUM W/ MANTLE SOURCE

EVOLVED = NOT IN EQ W/ SOURCE DUE TO FRACTIONAL CRYSTALLIZATION OR CONTAMINATION

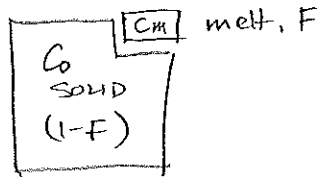
How do we get T_p ?

F , T , P of MELTING?

F CAN BE DETERMINED FROM INCOMPATIBLE TRACE ELEMENTS

SIMPLE APPROACH -

BATCH EQUILIBRIUM MELTING



PARTITION COEFFICIENT

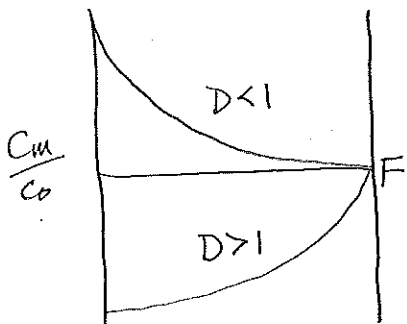
$$D = \frac{C_{\text{SOLID}}}{C_{\text{MELT}}}$$

- > 1 COMPATIBLE
- < 1 INCOMPATIBLE

$$C_0 = C_m F + C_s (1-F)$$

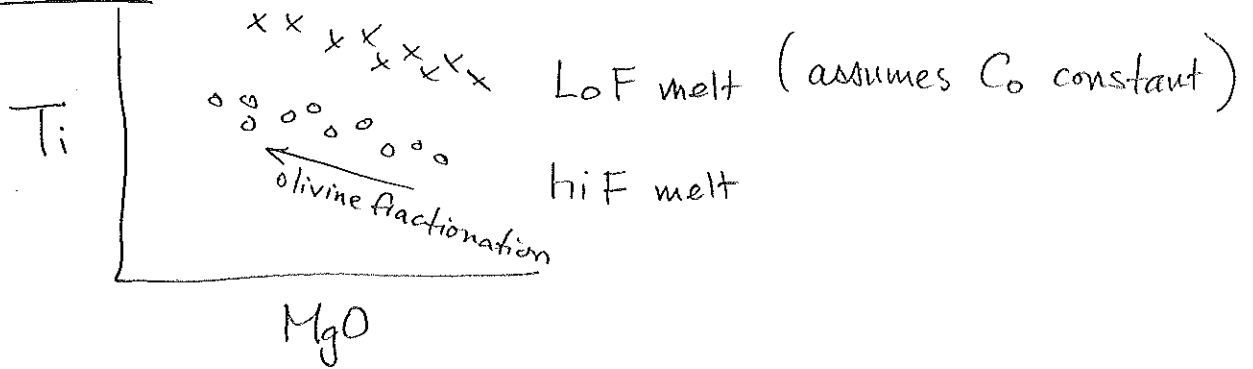
$$\frac{C_m}{C_0} = \frac{1}{D + F(1-D)}$$

← melt enrichment factor

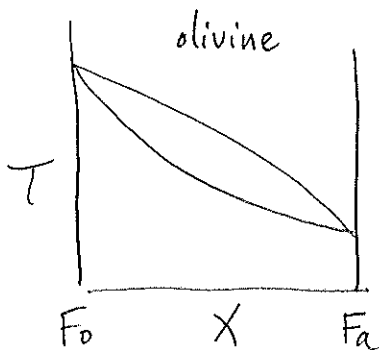


general rules: IF $D=0$ $\frac{C_m}{C_0} = \frac{1}{F}$
 IF $F=0$ $\frac{C_m}{C_0} = \frac{1}{D}$

application

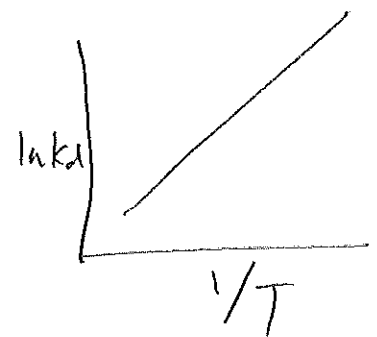


TEMPERATURE - THERMOMETER



$$K_d = \frac{MgO_{olivine}}{MgO_{Liquid}}$$

$$\ln K_d = \frac{a}{T} + b$$

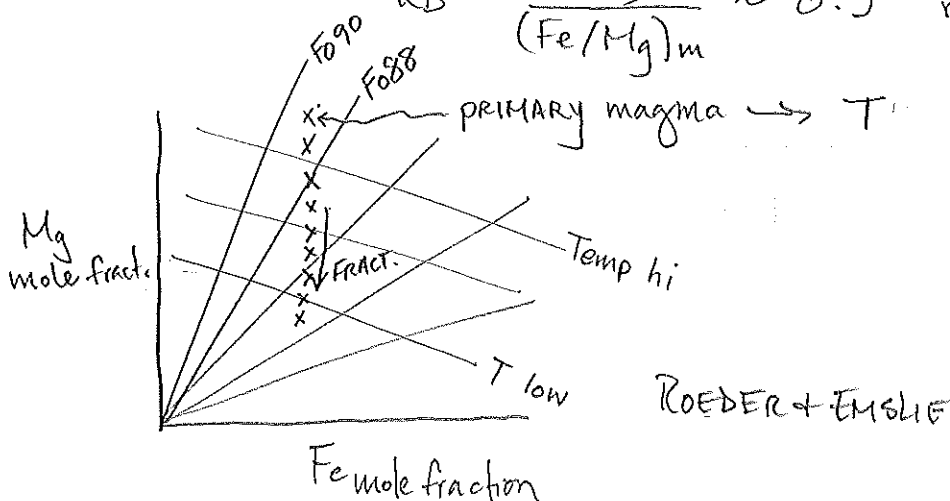


FROM MgO content of melt, you can get T.
 * High MgO → hi T (as long as olivine saturated)

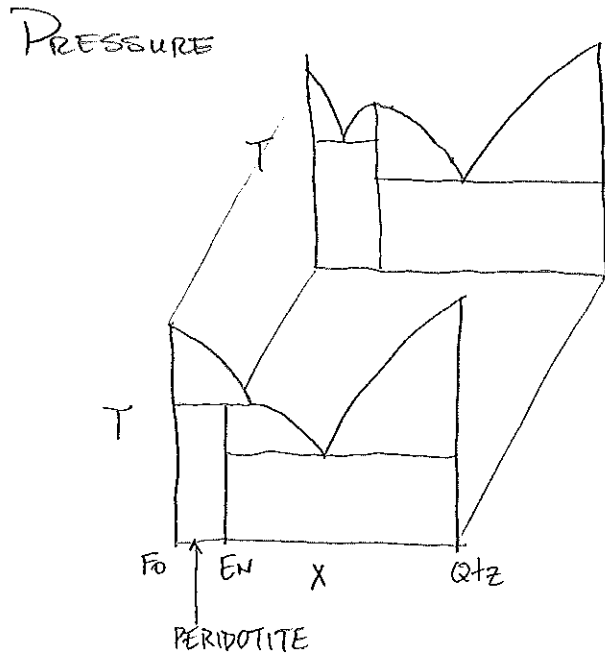
... (U) DETAIL



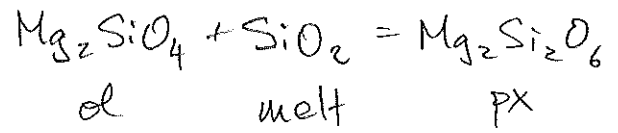
$$K_D = \frac{(Fe/Mg)_{ol}}{(Fe/Mg)_m} \sim 0.3 \text{ relatively constant with } T, P$$



ROEDER + EMSLIE



A PRESSURE-SENSITIVE RXN



$$K = \frac{a_{\text{PX}}^{\text{EN}}}{a_{\text{Ol}}^{\text{Fo}} a_{\text{melt}}^{\text{SiO}_2}}$$

$$\frac{a_{\text{PX}}^{\text{EN}}}{a_{\text{Ol}}^{\text{Fo}}} \sim \text{constant}$$

$$K \sim \frac{1}{a_{\text{melt}}^{\text{SiO}_2}} \quad \text{OLIV+OPX SYSTEM BUFFERS SiO}_2$$

RULES OF THUMB

* Low SiO₂ → hi P (assuming you have a primary magma)

CAVEATS/COMPLICATIONS

- MANTLE SOURCE VARIABLE IN TERMS OF MAJOR COMPOSITION
- PRIMARY MAGMA NOT KNOWN
- MAGMAS ARE AGGREGATES OR REACTED PRODUCTS.